

Impact of off-diagonal cross-shell interaction on ^{14}C *

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Abstract: A shell-model investigation is performed to show the impact on the structure of ^{14}C from the off-diagonal cross-shell interaction, $\langle pp|V|sdsd\rangle$, which represents the mixing between the 0 and $2\hbar\omega$ configurations in the *psd* model space. The observed levels of the positive states in ^{14}C can be nicely described in $0-4\hbar\omega$ or a larger model space through the well defined Hamiltonians, YSOX and WBP, with a reduction of the strength of the $\langle pp|V|sdsd\rangle$ interaction in the latter. The observed $B(\text{GT})$ values for ^{14}C can be generally described by YSOX, while WBP and their modifications of the $\langle pp|V|sdsd\rangle$ interaction fail for some values. Further investigation shows the effect of such interactions on the configuration mixing and occupancy. The present work shows examples of how the off-diagonal cross-shell interaction strongly drives the nuclear structure.

Keywords: shell model, off-diagonal cross-shell interaction, *psd* region, ^{14}C , levels, transition rates

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1 Introduction

The investigation of the nuclear interaction is of great importance in nuclear physics. Generally speaking, two approaches are used to study the nuclear interaction, starting from the realistic nucleon-nucleon (NN) force and from the nuclear data including binding energies and levels. The two approaches are used together and compared with each other in various nuclear models, such as the nuclear shell model [1].

To use a realistic NN force in a shell-model investigation, two problems need to be overcome: the strong short-range repulsion and the truncated model space [2, 3]. In the construction of the effective shell-model Hamiltonian, the latter problem is normally solved through many-body perturbation theory, which has certain difficulties in dealing with the cross-shell interaction (the interaction between different major oscillator shells) [2]. Recently, the extended Kuo-Krenciglowa (EKK) method has been suggested to derive the effective interaction among several oscillator shells [4, 5]. Its applications in the *sdpf* region show a nice agreement with the observed data, focusing on the binding energies, the levels of two-nucleon pairs, and the energies of the first 2^+ and 4^+ states in the even-even nuclei [6, 7].

Many effective Hamiltonians are well-defined considering the observed binding energies and levels, while

some of them have a realistic basis. Some examples are CK for the *p* shell [8], the USD family for the *sd* shell [9, 10], GXPF1 for the *pf* shell [11], MK [12], WBT [13], WBP [13] and SFO [14] for *psd* shells, and SDPF-M [15] for *sdpf* shells.

There are two types of cross-shell interaction in the two-body part of an effective Hamiltonian constructed for two major oscillator shells, which are $\langle N, N+1|V|N, N+1\rangle$ and $\langle N, N|V|N+1, N+1\rangle$, where N and $N+1$ are one and its next major oscillator shells, respectively. The first type, especially its diagonal part, is very important for the investigation of neutron-rich nuclei, where the protons are in the N shell and the neutrons are in the $N+1$ ($N+2$ for some extreme cases) shell. Its strength can be determined by considering the observed data of those nuclei. However, much less is known for the second type, which is purely off-diagonal, corresponding to the mixing between $n\hbar\omega$ and $(n+2)\hbar\omega$ configurations, where n means the number of nucleons excited to the next major shell. Some early effective Hamiltonians were constructed without consideration of the mixing between the 0 and $2\hbar\omega$ configurations, including MK, WBP, and WBT. The $\langle psd|V|psd\rangle$ interaction in WBP is obtained from a potential fitted to the observed data [13]. The $\langle pp|V|sdsd\rangle$ interactions of WBP and WBT are calculated from the same potential without considering its effect on the nuclear structure [13]. Later suggested

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effective Hamiltonians SFO, constructed in the $0-3\hbar\omega$ model space, did not focus on the properties of this interaction, but on the spin properties of the p shell nuclei [14].

It is seen that the off-diagonal cross-shell interaction, connecting the $n\hbar\omega$ and $(n+2)\hbar\omega$ configurations, is not well investigated in either realistic or phenomenological investigations. One reason is that its effect on nuclear structure is “hidden”, which means it is hard to show in an easily understandable scheme, such as the effect of the monopole interaction on the binding energies [16] and the shell structures [17–19]. Recently, such multipole correlations between normal and intruder configurations have been investigated in the neutron-rich nuclei around $N=20$ and 28 in the $sdpf$ model space [20]. The importance of the multi- $\hbar\omega$ configuration mixing has been investigated through the $Sp(3, \mathbb{R})$ shell model [21]. Clear symplectic symmetry in low-lying states of ^{12}C and ^{16}O shows that their NCSM wave functions can be typically projected to a few of the most deformed symplectic basis states at the level of 85%–90% [22]. The *ab initio* symmetry-adapted NCSM (SA-NCSM) results show that multi- $\hbar\omega$ configuration mixing is important for the description of the collective modes in light nuclei, such as ^6Li , ^6He , and ^8Be [23, 24]. The no-core symplectic shell model (NCSpm) is used to investigate multi- $\hbar\omega$ configuration mixing in α -clustering substructures in the low-lying states of ^{12}C [25] and in the ground state rotational bands of $^{20,22,24}\text{Ne}$, ^{20}O , $^{20,22}\text{Mg}$ and ^{24}Si [26].

The effective Hamiltonian YSOX for the psd region includes the effect of the off-diagonal cross-shell interaction based on the binding energies of B, C, N, and O isotopes from the stability line to the neutron drip line [27]. The results show that the strength of the central part of the $\langle pp|V|sdsd\rangle$ interaction is weaker than that of the $\langle psd|V|psd\rangle$ interaction. The effect of the strength of the $\langle pp|V|sdsd\rangle$ interaction on the low-lying levels of ^{10}B and ^{17}C was also presented. But the absolute values of the levels of these two nuclei do not vary significantly with the strength of the $\langle pp|V|sdsd\rangle$ interaction [27]. It is necessary to find more solid evidence on the effect of such interactions on the nuclear structure. For example, the levels and transition rates of some states, which are dramatically influenced by the $\langle pp|V|sdsd\rangle$ interaction, can be well described through YSOX, and other Hamiltonians with modifications of the strength of this interaction.

A good candidate for the above considerations is ^{14}C , one of the best known isotopes. The long lifetime of ^{14}C is a long-standing problem for theoretical models. It can be understood by the cancellation of the transition matrix elements between two main components in the p shell [1]. A few microscopic approaches based on the NN (and NNN) force have been studied to investigate

the origin of the extreme small $B(\text{GT})$ value between the ground states of ^{14}N and ^{14}C [28–32]. It should be noted that the energies of the first few positive states of ^{14}C and the $B(\text{GT})$ values for these states cannot be well described in the above microscopic investigations, the antisymmetrized molecular dynamics (AMD) method [33], or the shell model in the $0-2\hbar\omega$ model space with existing Hamiltonians, such as WBP, SFO, and YSOX [27]. The level of ^{14}C are normally excluded in the construction of a Hamiltonian for the global psd region.

The sd shell configurations are obviously important for the levels of ^{14}C . The valence protons and neutrons in ^{14}C fully occupy the $Z=6$ sub-shell and $N=8$ major shell, respectively. The single particle states of ^{13}C [34] indicate that the excitation energies inside the p shell have the same magnitude as the two-nucleon excitation energies from the p to the sd shell. Our recent conference proceeding [35] shows that the first few positive states are some of the mixing between $0\hbar\omega$ ($\sim 80\%$) and $2\hbar\omega$ ($\sim 20\%$) configurations, 0_1^+ , 2_1^+ , and 1_1^+ states, some of the almost pure $2\hbar\omega$ configuration, 0_2^+ and 2_2^+ states, and the pure $2\hbar\omega$ configuration, the 4_1^+ state, which is not possible to be coupled inside the p shell. The excitation energies of the 0_2^+ , 0_3^+ , and 4_1^+ states can be well described in a simple model of $(sd)^2$ states [36, 37].

In this paper, the structure of ^{14}C is investigated in the framework of the shell model up to the $6\hbar\omega$ excitation. It is shown that both the strength of the $\langle pp|V|sdsd\rangle$ interaction and the inclusion of the $4\hbar\omega$ configuration are crucial to reproduce the observed data of ^{14}C . The details of the Hamiltonian used in the present work is briefly introduced in Section 2. Levels and transition rates are discussed in Section 3 and 4, respectively. Some further discussions are presented in Section 5.

2 Hamiltonian

The nuclear shell model is widely used to investigate the structure of light and medium mass nuclei [38, 39]. In the psd region, the Hamiltonians MK, WBT, and WBP are successful. Both the $\langle pp|V|sdsd\rangle$ and $\langle pp|V|sdsd\rangle$ interactions in WBP are calculated through the same potential, which is convenient for discussing the different strengths in the present study. Thus WBP and its modifications on $\langle pp|V|sdsd\rangle$ interaction are considered in the following discussions.

The recently suggested Hamiltonian YSOX is also used in the present investigation. The $\langle pp|V|pp\rangle$ and $\langle sdsd|V|sdsd\rangle$ parts of YSOX are from the corresponding parts of SFO and SDPF-M, respectively [27]. The two types of cross-shell interaction, $\langle psd|V|psd\rangle$ and $\langle pp|V|sdsd\rangle$, are calculated through V_{MU} [19] plus spin-orbit force from M3Y [40]. V_{MU} is the monopole based universal interaction including a Gaussian type central

force and a $\pi+\rho$ bare tensor force, which assumes that the renormalization effect is mostly included in the central force [19, 41]. The validity of taking V_{MU} plus a spin-orbit term as the cross-shell interaction in the shell model is examined in various works for regions besides the psd region, such as the $sdpf$ region [42, 43] and the $pfsg$ region [44]. Such nuclear force is used to estimate the reduction effect caused by the weakly bound proton $1s_{1/2}$ orbit [45] and taken as the cross-shell interaction between two major shells, the $Z=28-50$ and $N=82-126$ shells [46]. The first $19/2^-$ state in ^{129}Pd is predicted to be a possible neutron-decaying isomer [46].

Among the above applications of V_{MU} , the renormalization effect is assumed to be contributed mostly by the central part. The tensor and spin-orbit parts are calculated through unchanged strength for both $\langle psd|V|psd\rangle$ and $\langle pp|V|sdsd\rangle$ interactions. The strengths of the central part of $\langle psd|V|psd\rangle$ and $\langle pp|V|sdsd\rangle$ interactions in YSOX are 0.85 and 0.55 of the original one, respectively [27]. The much weaker strength of the $\langle pp|V|sdsd\rangle$ interaction gives nice descriptions of the binding energies of the B, C, N, and O isotopes. The effect of the strength of such a central force is shown for low lying levels of ^{10}B and ^{17}C , but the change is not remarkable (≤ 0.5 MeV), when the strength varies from 0.55 to 0.85 (YSOX+) or 0.25 (YSOX-) of the original value [27]. The Hamiltonian YSOX+, with the same strength for both types of cross-shell interaction, is also used in the present work for comparison.

The two-body matrix elements (TBME) of WBP are compared in Ref. [27] with those of YSOX for central, spin-orbit, tensor and total interactions through the spin-tensor decomposition method [47]. It is found that the two types of cross-shell interaction are quite different between YSOX and WBP, except for the spin-orbit one. A modified version WBP- is introduced by multiplying all central TBME of the $\langle pp|V|sdsd\rangle$ interaction in WBP by a factor 0.6, which is similar to the reduction in YSOX.

The Hamiltonians WBP, WBP-, YSOX+, and YSOX are used in the following discussions for comparisons with each other. The shell-model calculations are performed with newly developed code KSHELL [48]. The center-of-mass (c.m.) correction is needed for multi-shell calculations. The standard method suggested by Gloeckner and Lawson [49] is used for the c.m. correction. They defined $H' = H_{\text{SM}} + \beta H_{\text{c.m.}}$, where H_{SM} and $H_{\text{c.m.}}$ are shell-model and c.m. Hamiltonians, respectively. In the present study, $\beta=10$ is adopted.

3 Levels of ^{14}C

The first few positive states of ^{14}C are not well described through various models because of the large shell gaps for both protons and neutrons. Some results obtained from NCSM [28] and AMD [33] are compared with those from the shell model in Fig. 1. None of the previous works and the shell-model results up to the $2\hbar\omega$ model space can reproduce the correct order of the first few positive states. For example, the 2^+ state is calculated to be the first excited positive state instead of the observed 0^+ state. It should be noted that both YSOX and WBP can give nice description of the levels, moments, and transition rates of the nearby nuclei. When the model space is up to $4\hbar\omega$ in the shell-model calculations, significant changes in the levels are found compared with those from the $0-2\hbar\omega$ model space. The excitation energies of $0_{2,3}^+$, $2_{2,3}^+$, and 4_1^+ are dramatically lower due to the inclusion of the $4\hbar\omega$ model space, while those of the 2_1^+ and 1_1^+ states vary little. The first five and the last two states are dominated by the $2\hbar\omega$ and $0\hbar\omega$ configurations, respectively. Such results agree with the analysis of $0_{2,3}^+$ and 4_1^+ states based on the $(sd)^2$ configuration [36, 37], $0_{2,3}^+$ and $2_{2,3}^+$ states based on the AMD method [33]. It should be mentioned that the 2_1^+ state obtained from YSOX in $0-4\hbar\omega$ model space has very strong mixing between the $0\hbar\omega$ and $2\hbar\omega$ configurations, which will be discussed in Section 5.

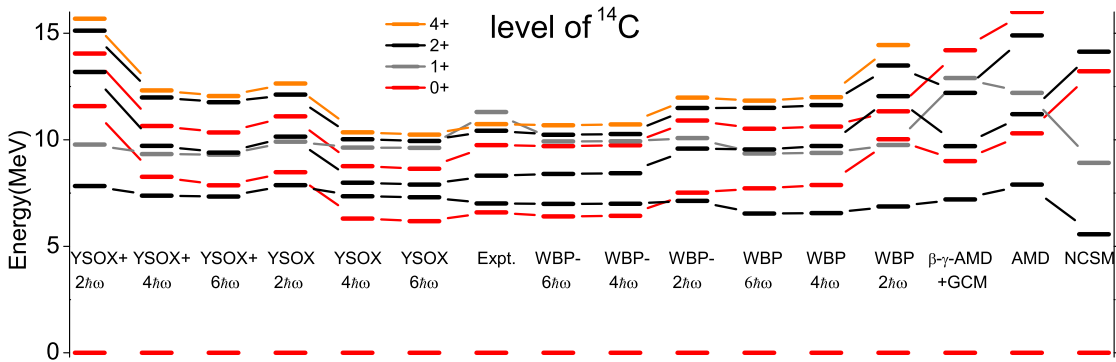


Fig. 1. (color online) Comparison of the energies of the positive states in ^{14}C between the observed data and the results of various models. Observed data, AMD, and NCSM results are taken from Refs. [34], [33], and [28], respectively

It is seen that the inclusion of the $4\hbar\omega$ model space is not enough to obtain a nice description of the energies of positive states in ^{14}C through the Hamiltonians, YSOX+ and well defined WBP. Both of them have the same strength in the two types of cross-shell interaction, $\langle psd|V|psd\rangle$ and $\langle pp|V|sdsd\rangle$. The Hamiltonians, YSOX and WBP-, can describe these energies well with just one modification, weakening of the $\langle pp|V|sdsd\rangle$ interaction. All shell-model results for the energy of the 1_1^+ state are lower than the observed value. This state is dominated by the excitation inside the p shell, rarely influenced by the higher $\hbar\omega$ excitation and the $\langle pp|V|sdsd\rangle$ interaction. Its excitation energy is not discussed further in the present work.

Figure 2 presents the binding energies of each of the states, giving a clear view of how the $\langle pp|V|sdsd\rangle$ interaction and the $4\hbar\omega$ configuration drive the evolution of the energies. It is seen that the binding energies of the $0\hbar\omega$ dominated states, 0_1^+ , 2_1^+ , and 1_1^+ , rise due to the weakening of the $\langle pp|V|sdsd\rangle$ interaction, while other states remain almost unchanged. The reason is that the $0\hbar\omega$ dominated states include certain percentages of the $2\hbar\omega$ configuration, while the $2\hbar\omega$ dominated states include rather few percentages of the $0\hbar\omega$ configuration. The energies contributed by the mixing between 0 and $2\hbar\omega$ configurations are weaker when the strength of the $\langle pp|V|sdsd\rangle$ interaction is reduced.

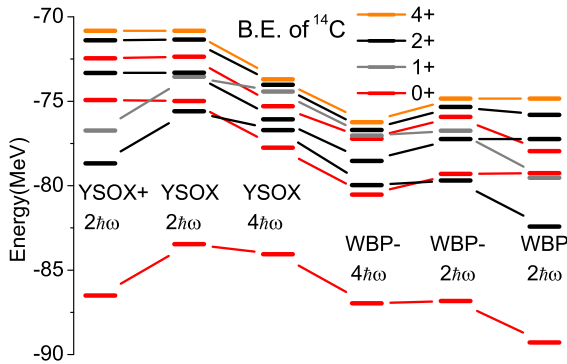


Fig. 2. (color online) Comparison of the binding energies of the positive states in ^{14}C among various shell-model calculations.

The inclusion of the $4\hbar\omega$ configuration leads to a different effect. The wave function of ^{14}C up to the $4\hbar\omega$ excitation is simply written as, $a|(sd)^0 + b|(sd)^2 + c|(sd)^4$, with the three terms corresponding to the 0, 2, and $4\hbar\omega$ configurations, respectively. The cross-shell interaction, $\langle pp|V|sdsd\rangle$, connects the first two and the last two configurations, but not the first and last one, because of its two-body nature. Thus, the inclusion of the $4\hbar\omega$ configuration does not show a significant effect on the $0\hbar\omega$ dominated states, but strongly affects the $2\hbar\omega$ dominated

states. The further inclusion of the $6\hbar\omega$ configuration results in rather small changes in level, because these states are not dominated by the $4\hbar\omega$ configuration.

The diagonal terms of the interaction, especially single particle energies, surely affect the levels of ^{14}C . Figure 3 shows the effect of the single particle energies on the levels of ^{14}C . The Hamiltonians YSOX_{spe}^+ and YSOX_{spe} are modified versions of YSOX+ and YSOX by reducing the gap between the p and sd shells by 1.0 MeV. The results indicate that the effect of reducing the gap is similar to the reduction of the $\langle pp|V|sdsd\rangle$ interaction and the increment of the model space from $2\hbar\omega$ to $4\hbar\omega$, by comparing among the levels from YSOX_{spe}^+ in $4\hbar\omega$, YSOX_{spe} in $2\hbar\omega$, and YSOX in $4\hbar\omega$. For example, seen from the YSOX_{spe}^+ and YSOX results, the reduced gap shows a similar effect to the reduction of the $\langle pp|V|sdsd\rangle$ interaction, which indicates that the latter reduction actually attracts the 0 and $2\hbar\omega$ configurations and increases the mixing between them. More details will be discussed in Section 5. The WBP results also show a similar effect by reducing the same gap by 0.9 MeV. Although the effect from the reduction on the gap is presented here, the reduction is not expected to be needed for YSOX in a real case, because the gap and the strength of the cross-shell interaction are simultaneously fixed to the single particle levels of ^{17}O , ^{15}C , ^{13}C , and other nuclei in the construction of the Hamiltonian YSOX. If the gap is changed, the strength of the cross-shell interaction also needs to be changed, which may give a worse description of global properties in nearby nuclei.

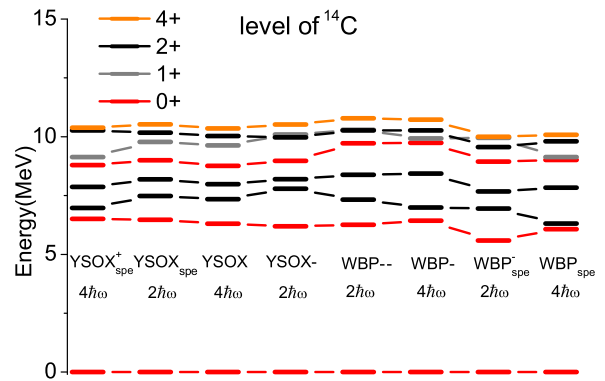


Fig. 3. (color online) Comparison of the energies of the positive states in ^{14}C among various shell-model calculations.

In a phenomenological view, the effect of the $\langle pp|V|sdsd\rangle$ interaction and the inclusion of the $4\hbar\omega$ configuration can be partially replaced by each other when concentrating on the levels of ^{14}C . Figure 3 shows similar results for the $4\hbar\omega$ calculations and for results with a further reduction in the central force in the $\langle pp|V|sdsd\rangle$ interaction. The Hamiltonians $\text{YSOX}-$ and $\text{WBP}-$ mean the strengths of the central force are reduced to

0.25 and 0.3 of their original values in V_{MU} and WBP, respectively. The results of YSOX and WBP– up to the $2\hbar\omega$ model space are similar to those of YSOX+ and WBP up to the $4\hbar\omega$ model space in Fig. 1. This similarity is not found, however, for YSOX and WBP– up to the $4\hbar\omega$ model space compared to YSOX+ and WBP up to the $6\hbar\omega$ model space. This means that the reduction of the $\langle pp|V|sdsd\rangle$ interaction cannot be fully replaced by the increment of the $n\hbar\omega$ excitation. It should be also noted that the strength of such a central force in YSOX is considered through the binding energies for all B, C, N, and O isotopes. A much larger or smaller value of the strength is not suitable for these binding energies [27].

In general, the energies of the first few positive states in ^{14}C can be well reproduced in the 0– $4\hbar\omega$ model space through the well defined Hamiltonians, YSOX and WBP, with a modification of the latter. Such a modification does not change the nice description of WBP in its original model space, apart from the mixing between the 0 and $2\hbar\omega$ states. Although some modifications shown in Fig. 3 similarly describe the levels of ^{14}C , the importance of the $\langle pp|V|sdsd\rangle$ interaction seems not to be substituted by other effects considering the global description of the nearby nuclei. The inclusion of the $4\hbar\omega$ in the present investigation is reasonable because of the existence of states dominated by the $2\hbar\omega$ configuration.

4 Transition rates

Besides the energies, the transition rates also show the effect of the $\langle pp|V|sdsd\rangle$ interaction. Figure 4 presents the $B(\text{GT})$ transition rates from the ground state of ^{14}N to the $0_{1,2}^+$, $2_{1,2,3}^+$, and 1_1^+ states of ^{14}C . Two quenching factors 0.72 and 0.64 are obtained to reproduce the observed $B(\text{GT})$ values among the nuclei around ^{14}C for YSOX and WBP, respectively [27]. The $B(\text{GT})$ values from the shell-model calculations in Fig. 4 are presented with these quenching factors, 0.72 for YSOX and YSOX+, 0.64 for WBP– and WBP, respectively.

The observed $B(\text{GT})$ values for the 0_1^+ and 2_1^+ states of ^{14}C are overestimated by all theoretical results, as shown in Fig. 4. The results for the former are not very clear in the figure because of its rather small absolute value. Such a small $B(\text{GT})$ value corresponds to the long lifetime of ^{14}C . It is shown that the value is very sensitive to the model space and the strength of the spin-orbit and tensor force [50]. Several microscopic NCSM investigations have been performed to investigate the origin of the small value [28–31]. NCSM with chiral $NN + NNN$ interactions can explain the rather small transition rate [51]. In general, YSOX and YSOX+ give smaller $B(\text{GT})$ values for 0_1^+ state compared with WBP– and WBP.

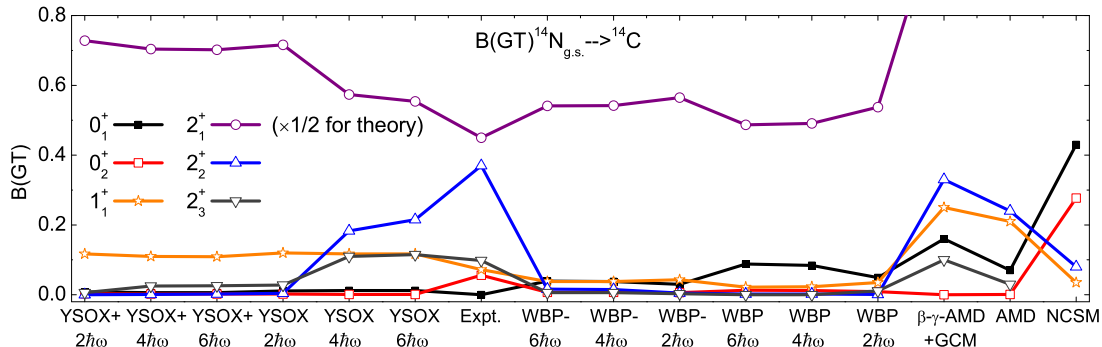


Fig. 4. (color online) Comparison of the $B(\text{GT})(^{14}\text{N}_{g.s.} \rightarrow ^{14}\text{C}_{0^+,1^+,2^+})$ between the observed data and the results of various models. Observed data, AMD, and NCSM results are taken from Refs. [52], [33], and [28], respectively.

The $B(\text{GT})$ values for the 2_1^+ state from the shell model are systematically slightly larger than twice the observed value. Certain deficiencies may exist in the descriptions of the 2_1^+ state of ^{14}C and/or the 1_1^+ state of ^{14}N . The NCSM [28] and AMD [33] $B(\text{GT})$ values for 2_1^+ state are around five times as high as the observed data, beyond the range of Fig. 4. The $B(\text{GT})$ value for 1_1^+ state is generally well described by the shell model and NCSM.

The $B(\text{GT})$ values for $2\hbar\omega$ dominated $2_{2,3}^+$ states cannot be reproduced except by the AMD method and the

Hamiltonian YSOX up to 4 and $6\hbar\omega$ model spaces. The $0_{2,3}^+$ and $2_{2,3}^+$ states include cluster correlations resulting from the mixing of higher shell components in AMD calculations, which can be described within the $6\hbar\omega$ model space [33]. The present shell-model study agrees with such a statement, because few differences are found in the level and $B(\text{GT})$ values of these states between the 4 and $6\hbar\omega$ results. But it is clearly seen that the inclusion of the higher shell components is not enough for the shell model to reproduce the $B(\text{GT})$ values for the $2_{2,3}^+$ states. Only the Hamiltonian YSOX can give good

descriptions of these two $B(\text{GT})$ values, while all others fail, including YSOX+, WBP, and WBP-. The complicated correlation can be described by the combination of the higher excitation to sd shell and the weakening of the $\langle pp|V|sdsd\rangle$ interaction based on V_{MU} plus spin-orbit

force. It is not easy to fully understand such effects, because it is difficult to know how off-diagonal interaction drives the structure of the nuclei. Some further discussion is given in the next section.

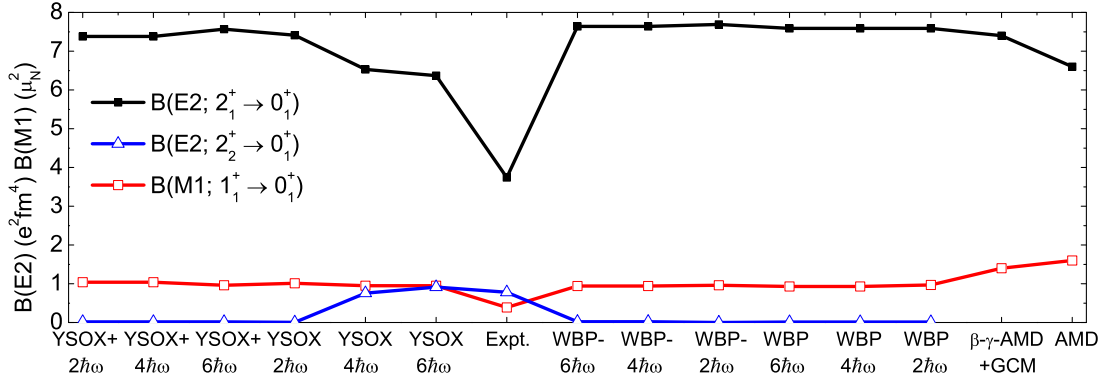


Fig. 5. (color online) Comparison of the $B(\text{E}2)(2_{1,2}^+ \rightarrow 0_1^+)$ and $B(\text{M}1)(1_1^+ \rightarrow 0_1^+)$ in ^{14}C between the observed data and the results of various models. Observed data and AMD results are taken from Refs. [53, 54] and [33], respectively.

Similar to the $B(\text{GT})$ value for the 2_1^+ state, the $B(\text{E}2)$ value for the same state and the $B(\text{M}1)$ value for 1_1^+ state are also overestimated by all theoretical approaches, as shown in Fig. 5. The effective charges $e_p=1.27$, $e_n=0.23$, the effective g factors $\delta g_{\pi,\nu}^{(l)}=\pm 0.1 \mu_N$ and $g_s^{(\text{eff})}/g_s=0.95$ are used in the present calculations, which are obtained through the systematic trends of the electromagnetic properties of B, C, N, and O isotopes in Ref. [27]. Very few discussions are found on the $B(\text{E}2)$ value for the 2_2^+ state, which is reported in Ref. [55]. This value is well reproduced by the results of YSOX in both the $4\hbar\omega$ and $6\hbar\omega$ model spaces, while all other calculations fail. In general, the YSOX results in $4\hbar\omega$ and $6\hbar\omega$ model spaces give better descriptions than other calculations of these transitions.

5 Further discussion

It is of great importance to know why YSOX gives a better description than WBP- for the $B(\text{GT})$ and $B(\text{E}2)$ values, while both of them can reproduce the levels of ^{14}C . YSOX shows better performance than WBP in a global comparison of the $B(\text{GT})$ values among the nuclei around ^{14}C [27]. The TBME and the monopole terms of YSOX and WBP are compared in Ref. [27]. A similar comparison of the $\langle pp|V|sdsd\rangle$ part between YSOX and WBP- is presented in Fig. 6. Please note that the only difference between WBP and WBP- is the central part of the $\langle pp|V|sdsd\rangle$ interaction. The comparison of the central TBME of the $\langle pp|V|sdsd\rangle$ interaction between YSOX and WBP- shows more similarities than

that between YSOX and WBP, which indicates why the reduction of such a central force in WBP- can reproduce the levels of ^{14}C .

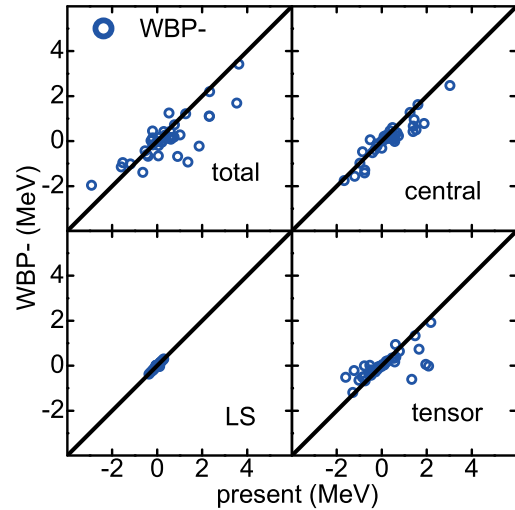


Fig. 6. (color online) Comparison of the TBME of $\langle pp|V|sdsd\rangle$ interaction between YSOX and WBP-.

YSOX and WBP- are similar in all $\langle sdsd|V|sdsd\rangle$ parts, the central and spin-orbit forces of the $\langle pp|V|pp\rangle$ part, and the spin-orbit force of the $\langle psd|V|psd\rangle$ and $\langle pp|V|sdsd\rangle$ parts, seen from Ref. [27]. Most differences come from the central force of the last two parts and

the tensor force of the last three parts, in total five components. The TBME of each of the five components in WBP– is replaced by the corresponding one in YSOX. However, none of the modified Hamiltonians can reproduce the $B(\text{GT})$ values for the $2_{2,3}^+$ states. This shows the complexity of the reason that such $B(\text{GT})$ values can be well described by YSOX, as there may be contributions from the combination of several of the five components.

Table 1. The transition matrix elements of the calculated $B(\text{GT})(^{14}\text{N}_{g.s.} \rightarrow ^{14}\text{C}_{2_{1,2,3}^+})$

state	Hamiltonian	model space	$M_{0p_{1/2} \rightarrow 0p_{3/2}}$	M_{other}
2_1^+	YSOX	$2\hbar\omega$	1.95	0.13
2_1^+	YSOX	$4\hbar\omega$	1.59	0.27
2_1^+	YSOX+	$4\hbar\omega$	1.80	0.26
2_1^+	WBP–	$4\hbar\omega$	1.73	0.07
2_2^+	YSOX	$2\hbar\omega$	0.31	-0.20
2_2^+	YSOX	$4\hbar\omega$	0.89	-0.15
2_2^+	YSOX+	$4\hbar\omega$	0.26	-0.21
2_2^+	WBP–	$4\hbar\omega$	0.28	-0.07
2_3^+	YSOX	$2\hbar\omega$	0.43	-0.13
2_3^+	YSOX	$4\hbar\omega$	0.72	-0.14
2_3^+	YSOX+	$4\hbar\omega$	0.46	-0.19
2_3^+	WBP–	$4\hbar\omega$	0.26	-0.13

This problem can be partly understood through detailed investigation of the transition matrix elements and the configurations. Table 1 presents the most important

transition matrix element $M_{0p_{1/2} \rightarrow 0p_{3/2}}$ of $B(\text{GT})(^{14}\text{N}_{g.s.} \rightarrow ^{14}\text{C}_{2_{1,2,3}^+})$ value. All calculations give very small transition matrix elements between other orbits for these three $B(\text{GT})$ values. All shell-model results in Fig. 4 give rather small $B(\text{GT})(^{14}\text{N}_{g.s.} \rightarrow ^{14}\text{C}_{2_{2,3}^+})$ except those from YSOX in 4 and $6\hbar\omega$ model spaces. The main difference comes from the $M_{0p_{1/2} \rightarrow 0p_{3/2}}$ term, which is much enhanced in the calculation from YSOX in the $4\hbar\omega$ model space, while the same transition matrix element in $B(\text{GT})(^{14}\text{N}_{g.s.} \rightarrow ^{14}\text{C}_{2_1^+})$ is smaller in the same set of calculations.

Table 2 shows the percentages of the 0, 2, and $4\hbar\omega$ configurations and the occupancies of each orbit in the $0_{1,2}^+$ and $2_{1,2,3}^+$ states of ^{14}C . Calculations from YSOX in the $4\hbar\omega$ model space give stronger mixing between 0 and $2\hbar\omega$ configurations in all three 2^+ states than those from all other calculations. A larger (smaller) $0\hbar\omega$ configuration in the $2_{2,3}^+$ (2_1^+) states leads to larger (smaller) $M_{0p_{1/2} \rightarrow 0p_{3/2}}$ terms in Table 1. From YSOX+ to YSOX, the reduction in the $\langle pp|V|sdsd\rangle$ interaction makes the 0 and $2\hbar\omega$ configurations more attractive in some states, such as the 2_1^+ state, but more repulsive in other states, such as the 0_1^+ state. When the interaction is reduced, the repulsive and attractive terms shown in Fig. 6 contribute to the attraction and repulsion of the 0 and $2\hbar\omega$ configurations, respectively. The spin-dependent nature of the nuclear interaction causes the differences in the 0_1^+ and 2_1^+ states.

Table 2. The configurations and the neutron occupancies of the $0_{1,2}^+$ and $2_{1,2,3}^+$ states in ^{14}C .

state	Hamiltonian	space	$0\hbar\omega(\%)$	$2\hbar\omega(\%)$	$4\hbar\omega(\%)$	$N_{p_{1/2}}$	$N_{p_{3/2}}$	$N_{d_{3/2}}$	$N_{d_{5/2}}$	$N_{s_{1/2}}$
0_1^+	YSOX	$2\hbar\omega$	83.39	16.61	-	1.87	3.91	0.07	0.13	0.02
0_1^+	YSOX	$4\hbar\omega$	77.95	20.90	1.15	1.81	3.87	0.09	0.20	0.04
0_1^+	YSOX+	$4\hbar\omega$	69.30	28.15	2.55	1.75	3.81	0.12	0.28	0.04
0_1^+	WBP–	$4\hbar\omega$	89.69	10.04	0.27	1.91	3.96	0.03	0.08	0.02
0_2^+	YSOX	$2\hbar\omega$	3.60	96.40	-	0.69	3.40	0.17	1.08	0.66
0_2^+	YSOX	$4\hbar\omega$	6.01	86.92	7.07	0.75	3.33	0.18	0.97	0.78
0_2^+	YSOX+	$4\hbar\omega$	6.12	82.28	11.60	0.77	3.29	0.16	0.78	1.01
0_2^+	WBP–	$4\hbar\omega$	2.01	94.78	3.21	0.71	3.33	0.16	1.25	0.56
2_1^+	YSOX	$2\hbar\omega$	75.13	24.87	-	1.78	3.87	0.09	0.23	0.04
2_1^+	YSOX	$4\hbar\omega$	49.12	47.47	3.41	1.42	3.67	0.14	0.63	0.15
2_1^+	YSOX+	$4\hbar\omega$	60.22	36.07	3.71	1.66	3.75	0.15	0.38	0.06
2_1^+	WBP–	$4\hbar\omega$	81.30	18.03	0.67	1.84	3.91	0.06	0.17	0.03
2_2^+	YSOX	$2\hbar\omega$	1.86	98.14	-	0.66	3.39	0.18	1.10	0.67
2_2^+	YSOX	$4\hbar\omega$	15.11	78.36	6.53	0.90	3.41	0.18	0.90	0.63
2_2^+	YSOX+	$4\hbar\omega$	1.26	86.72	12.02	0.70	3.26	0.20	1.13	0.71
2_2^+	WBP–	$4\hbar\omega$	2.21	94.09	3.70	0.71	3.33	0.15	1.25	0.56
2_3^+	YSOX	$2\hbar\omega$	3.92	96.08	-	0.81	3.29	0.12	1.50	0.28
2_3^+	YSOX	$4\hbar\omega$	10.41	83.11	6.48	0.92	3.27	0.13	1.45	0.23
2_3^+	YSOX+	$4\hbar\omega$	4.52	84.73	10.75	0.83	3.21	0.14	1.58	0.24
2_3^+	WBP–	$4\hbar\omega$	2.30	93.81	3.89	0.82	3.21	0.13	1.46	0.38

The sd shell neutrons in the 2_1^+ state mainly occupy the $0d_{5/2}$ orbit. Some important TBME, $\langle p_{1/2}p_{1/2}|V|d_{5/2}d_{5/2}\rangle$ and $\langle p_{3/2}p_{3/2}|V|d_{5/2}d_{5/2}\rangle$, are repulsive and contribute to the enhanced occupancy on the $0d_{5/2}$ orbit from YSOX+ to YSOX. Strong mixing between the 0 and $2\hbar\omega$ configurations in $2_{1,2}^+$ is also suggested, based on the analysis of inelastic pion scattering [56]. It should be noted that the phenomenological shell-model approaches normally give less multi- $\hbar\omega$ mixing compared with the NCSM based methods, because a phenomenological Hamiltonian is normally fitted with the assumption that $0\hbar\omega$ states are dominant states in most nuclei considered in the model space. For example, NCSM gives 56% and 51% $0\hbar\omega$ configurations for the ground states of ^{12}C and ^{16}O , respectively [22] and NCSpM gives around 65% $0\hbar\omega$ configurations for the former state [25], while YSOX in $4\hbar\omega$ model space gives 86% and 69% $0\hbar\omega$ configurations for these two states, respectively.

In light nuclei, α -cluster structures may occur. One famous example is the Hoyle state. It is shown that the Hoyle state demands 4– $14\hbar\omega$ states in a NCSpM study [25], which is difficult to describe in the present phenomenological approach (effective Hamiltonians normally cannot give the proper position of the Hoyle state). Because of the two extra neutrons, the low-lying states of ^{14}C should be dominated by nucleon(s) excitation, rather than the 3α structure of the Hoyle state. Thus the YSOX 0– $4\hbar\omega$ results are not much different from those with the 0– $6\hbar\omega$ model space. Although the 3α structure may be excluded in the low-lying states of ^{14}C , it is still difficult to identify whether there are α structure in these states in the present approach.

The effective single-particle energies [57] in these states can be considered with the occupancies obtained from the shell model, as they are shown for the neutron rich C, N, and O isotopes [58]. The differences in the occupancies presented in Table 2 among different sets of the calculations thus change the single-particle structures. The single-particle structure and the occupancy are normally dominated by the diagonal TBME, especially the monopole terms. The effect of the off-diagonal TBME on the occupancy is shown in the present work as a special example. It should be emphasized that the off-diagonal cross-shell interaction may have more effects on the configuration mixing beyond the configurations and occupancies, which demands further investigation.

The uncertainty of a theoretical model normally consists of two parts, statistical uncertainty from the parameters which are not well determined and systematic uncertainty from the deficiencies of the model. Recently, these two types of uncertainty have been analysed for the liquid drop model based on the uncertainty decomposition method [59]. The present work provides a preliminary analysis of the systematic uncertainty of the shell model, which comes from the constraints on the model space.

6 Summary

In summary, the present work has investigated the effect of the off-diagonal cross-shell interaction on the levels and the transition rates of ^{14}C . Based on two well defined Hamiltonians in the psd shell, WBP and YSOX, the observed excitation energies of the first few positive states in ^{14}C can be described well in the 0– $4\hbar\omega$ model space. A weaker strength of the $\langle pp|V|sdsd\rangle$ interaction is needed for WBP, while no changes are necessary for YSOX. It should be mentioned that the strength of the $\langle pp|V|sdsd\rangle$ interaction is not considered in the construction of WBP.

The $B(\text{GT})$ transition rates between the ground state of ^{14}N and the positive states of ^{14}C can be generally described by YSOX in the 0– $4\hbar\omega$ or a larger model space under the framework of the nuclear shell model, but other Hamiltonians fail. Although it is hard to fully understand such results, further comparisons of the transition matrix elements and the configurations show that part of the reason comes from the strong mixing between the 0 and $2\hbar\omega$ configurations in the 2^+ states of ^{14}C from YSOX results. The strong mixing is contributed by both the increment of the model space and the reduction of the off-diagonal cross-shell interaction. More effects may exist because of the complexity of the way that the off-diagonal TBME drives the nuclear structure.

The effect of the off-diagonal cross-shell interaction on the nuclear structure is not well investigated because it is less “visible”. The present work shows some examples of its effect on the structure of ^{14}C , while the Hamiltonians are constructed from the good global descriptions of nearby nuclei. It is of great importance to perform more investigations on this interaction based on both phenomenological and realistic approaches.

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